



CHALMERS
UNIVERSITY OF TECHNOLOGY



Kinetic modeling of argon-induced disruptions in ASDEX Upgrade

K. Insulander Björk¹

G. Papp², O. Embréus¹, L. Hesslow¹, T. Fülöp¹, O. Vallhagen¹, A. Lier², G. Pautasso², A. Bock², the ASDEX Upgrade Team^{2*} and the EUROfusion MST1 Team[†]

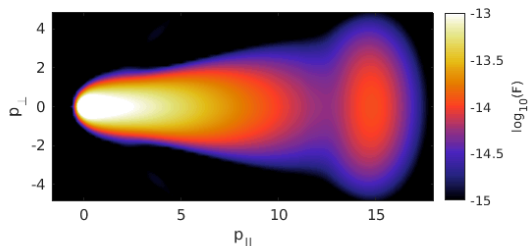
¹Chalmers University of Technology, Gothenburg, Sweden

²Max Planck Institute for Plasma Physics, Garching, Germany

* See author list of "H. Meyer *et al.* 2019 *Nucl. Fusion* **59** 112014"

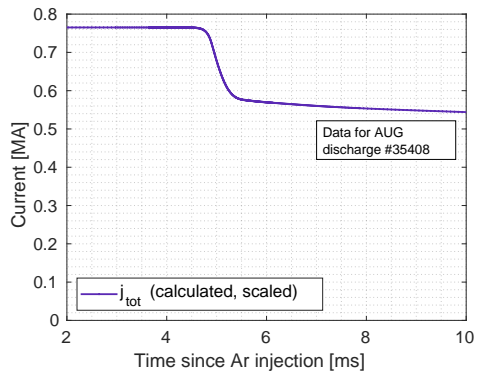
† See the author list of "B. Labit *et al.* 2019 *Nucl. Fusion* **59** 086020"



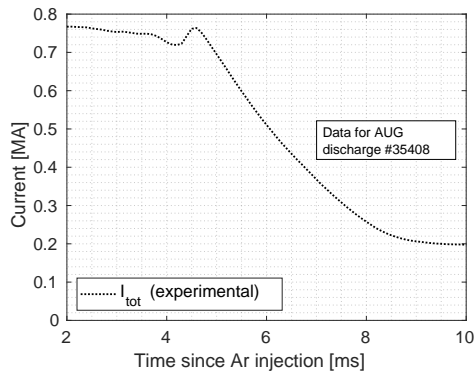


1. Argon induced disruptions in AUG
2. Temperature data
3. Density data
4. Current evolution
5. Distribution functions
6. RE generation rates by different mechanisms
7. Comparing model and experiment for 10 shots...
8. Summary

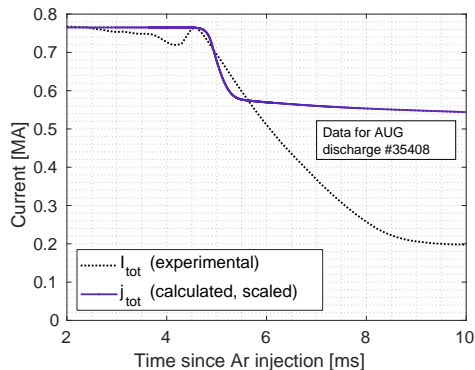
- Main take-home messages from REM 2019:



- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes

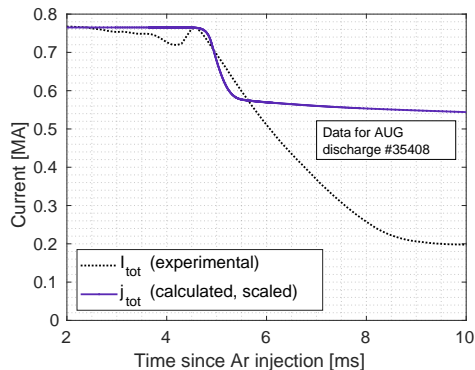


- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments

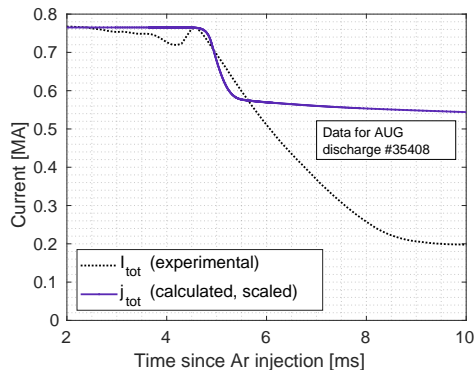


■ Main take-home messages from REM 2019:

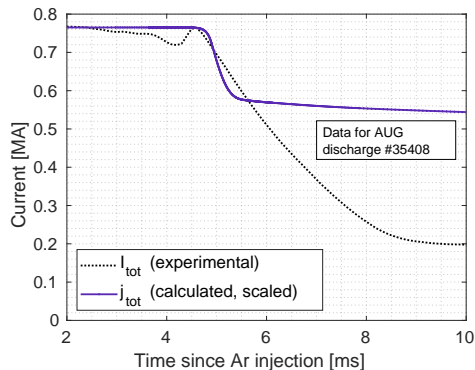
- ▶ Theorists make awesome codes
- ▶ Experimentalists make awesome experiments
- ▶ Fitting the former with the latter is not so awesome



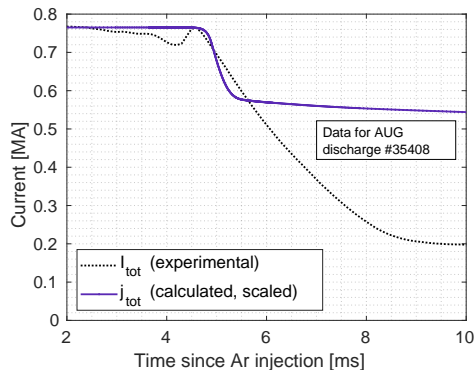
- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments
 - ▶ Fitting the former with the latter is not so awesome
- Main take-home messages from my stay at ASDEX-Upgrade:



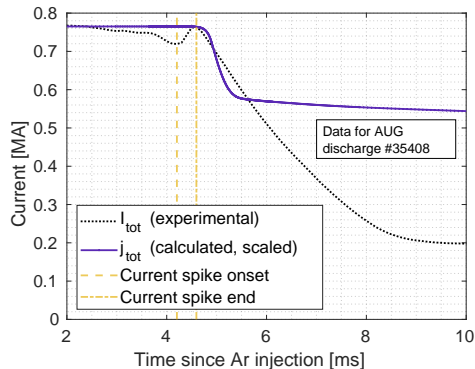
- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments
 - ▶ Fitting the former with the latter is not so awesome
- Main take-home messages from my stay at ASDEX-Upgrade:
 - ▶ Some things can be measured



- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments
 - ▶ Fitting the former with the latter is not so awesome
- Main take-home messages from my stay at ASDEX-Upgrade:
 - ▶ Some things can be measured
 - ▶ Other things can be estimated



- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments
 - ▶ Fitting the former with the latter is not so awesome
- Main take-home messages from my stay at ASDEX-Upgrade:
 - ▶ Some things can be measured
 - ▶ Other things can be estimated
 - ▶ Most things can only be guessed



- Main take-home messages from REM 2019:
 - ▶ Theorists make awesome codes
 - ▶ Experimentalists make awesome experiments
 - ▶ Fitting the former with the latter is not so awesome
- Main take-home messages from my stay at ASDEX-Upgrade:
 - ▶ Some things can be measured
 - ▶ Other things can be estimated
 - ▶ Most things can only be guessed
 - ▶ Especially so during MHD events...

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche
- Self-consistent electric field evolution

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche
- Self-consistent electric field evolution
- Collisions: relativistic test particle operator [2,3]

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

[2] L. Hesslow, et al.: *Journal of Plasma Physics*, **85**(6), 2019

[3] B. J. Braams, et al.: *Physics of Fluids B: Plasma Physics*, **1**(7):1355, 1989

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche
- Self-consistent electric field evolution
- Collisions: relativistic test particle operator [2,3]
- Screening of partially ionized impurities [4]

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

[2] L. Hesslow, et al.: *Journal of Plasma Physics*, **85**(6), 2019

[3] B. J. Braams, et al.: *Physics of Fluids B: Plasma Physics*, **1**(7):1355, 1989

[4] L. Hesslow, et al.: *Physical Review Letters*, **118**, 2017

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche
- Self-consistent electric field evolution
- Collisions: relativistic test particle operator [2,3]
- Screening of partially ionized impurities [4]
- Radiation losses (synchrotron and Bremsstrahlung)

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

[2] L. Hesslow, et al.: *Journal of Plasma Physics*, **85**(6), 2019

[3] B. J. Braams, et al.: *Physics of Fluids B: Plasma Physics*, **1**(7):1355, 1989

[4] L. Hesslow, et al.: *Physical Review Letters*, **118**, 2017

The relativistic finite-difference Fokker-Planck solver CODE [1] simulates electron dynamics in plasmas:

- 2D momentum space, 0D real space (present simulations: on-axis)
- Runaway electron generation: hot-tail, Dreicer and avalanche
- Self-consistent electric field evolution
- Collisions: relativistic test particle operator [2,3]
- Screening of partially ionized impurities [4]
- Radiation losses (synchrotron and Bremsstrahlung)
- No radial transport or instabilities.

[1] A. Stahl, et al.: *Nuclear Fusion*, **56**(11), 2016

[2] L. Hesslow, et al.: *Journal of Plasma Physics*, **85**(6), 2019

[3] B. J. Braams, et al.: *Physics of Fluids B: Plasma Physics*, **1**(7):1355, 1989

[4] L. Hesslow, et al.: *Physical Review Letters*, **118**, 2017

- Parameters that are (more or less) similar in all shots:
 - ▶ Magnetic field $B = 2.5$ T

- Parameters that are (more or less) similar in all shots:
 - ▶ Magnetic field $B = 2.5$ T
 - ▶ Major radius $R = 1.65$ m

- Parameters that are (more or less) similar in all shots:
 - ▶ Magnetic field $B = 2.5$ T
 - ▶ Major radius $R = 1.65$ m
 - ▶ Minor radius $a = 0.5$ m

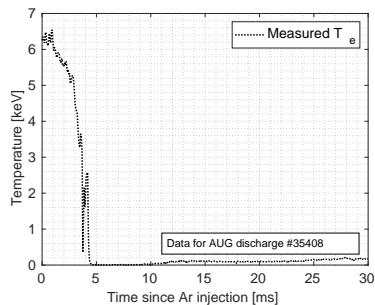
- Parameters that are (more or less) similar in all shots:
 - ▶ Magnetic field $B = 2.5$ T
 - ▶ Major radius $R = 1.65$ m
 - ▶ Minor radius $a = 0.5$ m
 - ▶ Initial current $I_0 = 0.7\text{--}0.8$ MA

| | Shot number |
|---|----------------|
| <ul style="list-style-type: none"> ■ Parameters that are (more or less) similar in all shots: <ul style="list-style-type: none"> ▶ Magnetic field $B = 2.5$ T ▶ Major radius $R = 1.65$ m ▶ Minor radius $a = 0.5$ m ▶ Initial current $I_0 = 0.7\text{--}0.8$ MA | # |
| <ul style="list-style-type: none"> ■ Parameters that vary between shots: | 35401 |
| | 34149 |
| | 34183 |
| | 34140 |
| | 34084 |
| | 35649 |
| | 35650 |
| | 35408 |
| | 33108 |
| | 31318 |

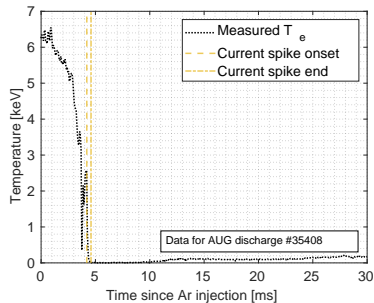
| ■ Parameters that are (more or less) similar in all shots: | Shot number | Injected argon [bar] |
|--|-------------|----------------------|
| ▶ Magnetic field $B = 2.5$ T | # | p_{Ar} |
| ▶ Major radius $R = 1.65$ m | 35401 | 0.15 |
| ▶ Minor radius $a = 0.5$ m | 34149 | 0.2 |
| ▶ Initial current $I_0 = 0.7\text{--}0.8$ MA | 34183 | 0.3072 |
| ■ Parameters that vary between shots: | 34140 | 0.31 |
| ▶ Injected argon quantity (1 bar, 0.1 l, 300 K $\iff 2.4 \cdot 10^{21}$ atoms) | 34084 | 0.33 |
| | 35649 | 0.39 |
| | 35650 | 0.4 |
| | 35408 | 0.5 |
| | 33108 | 0.73 |
| | 31318 | 0.9 |

| | Shot number | Injected argon [bar] | Initial density [m^{-3}] |
|---|-------------|----------------------|-------------------------------------|
| | # | p_{Ar} | n_0 |
| <ul style="list-style-type: none"> ■ Parameters that are (more or less) similar in all shots: <ul style="list-style-type: none"> ▶ Magnetic field $B = 2.5$ T ▶ Major radius $R = 1.65$ m ▶ Minor radius $a = 0.5$ m ▶ Initial current $I_0 = 0.7\text{--}0.8$ MA | 35401 | 0.15 | $2.6 \cdot 10^{19}$ |
| | 34149 | 0.2 | $3.0 \cdot 10^{19}$ |
| | 34183 | 0.3072 | $2.8 \cdot 10^{19}$ |
| | 34140 | 0.31 | $2.3 \cdot 10^{19}$ |
| <ul style="list-style-type: none"> ■ Parameters that vary between shots: <ul style="list-style-type: none"> ▶ Injected argon quantity (1 bar, 0.1 l, 300 K $\iff 2.4 \cdot 10^{21}$ atoms) ▶ Initial free electron density | 34084 | 0.33 | $3.0 \cdot 10^{19}$ |
| | 35649 | 0.39 | $2.6 \cdot 10^{19}$ |
| | 35650 | 0.4 | $2.4 \cdot 10^{19}$ |
| | 35408 | 0.5 | $2.4 \cdot 10^{19}$ |
| | 33108 | 0.73 | $3.1 \cdot 10^{19}$ |
| | 31318 | 0.9 | $2.2 \cdot 10^{19}$ |

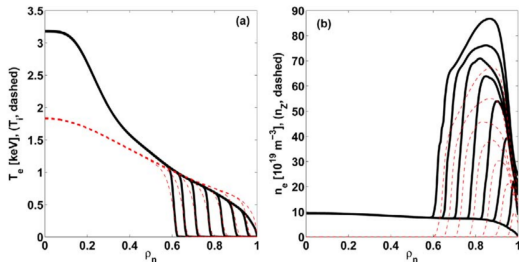
| | Shot number | Injected argon [bar] | Initial density [m^{-3}] | Initial temperature [keV] |
|---|-------------|----------------------|------------------------------|---------------------------|
| | # | p_{Ar} | n_0 | T_{e0} |
| <ul style="list-style-type: none"> ■ Parameters that are (more or less) similar in all shots: <ul style="list-style-type: none"> ▶ Magnetic field $B = 2.5$ T ▶ Major radius $R = 1.65$ m ▶ Minor radius $a = 0.5$ m ▶ Initial current $I_0 = 0.7-0.8$ MA | 35401 | 0.15 | $2.6 \cdot 10^{19}$ | 6.1 |
| | 34149 | 0.2 | $3.0 \cdot 10^{19}$ | 5.7 |
| | 34183 | 0.3072 | $2.8 \cdot 10^{19}$ | 5.5 |
| | 34140 | 0.31 | $2.3 \cdot 10^{19}$ | 5.8 |
| <ul style="list-style-type: none"> ■ Parameters that vary between shots: <ul style="list-style-type: none"> ▶ Injected argon quantity (1 bar, 0.1 l, 300 K $\iff 2.4 \cdot 10^{21}$ atoms) ▶ Initial free electron density ▶ Initial free electron temperature | 34084 | 0.33 | $3.0 \cdot 10^{19}$ | 4.3 |
| | 35649 | 0.39 | $2.6 \cdot 10^{19}$ | 6.2 |
| | 35650 | 0.4 | $2.4 \cdot 10^{19}$ | 5.3 |
| | 35408 | 0.5 | $2.4 \cdot 10^{19}$ | 6.0 |
| | 33108 | 0.73 | $3.1 \cdot 10^{19}$ | 7.2 |
| | 31318 | 0.9 | $2.2 \cdot 10^{19}$ | 11 |

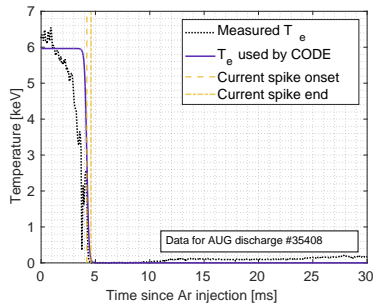


- Free electron temperature measured by ECE

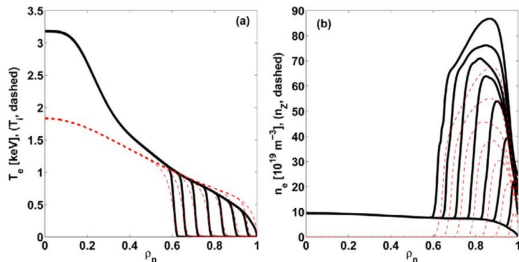


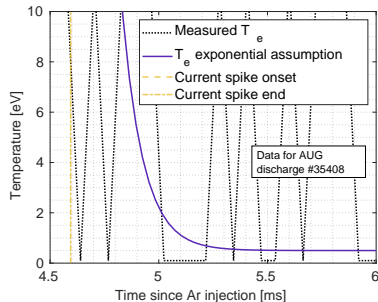
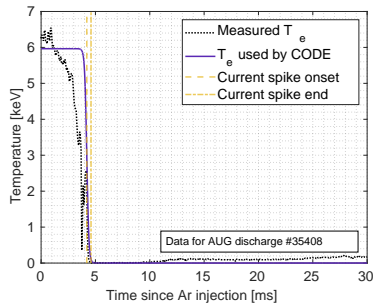
- Free electron temperature measured by ECE
- ECE blocked by high e^- densities on edge ≈ 1 ms after injection



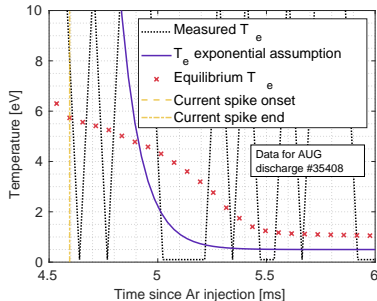
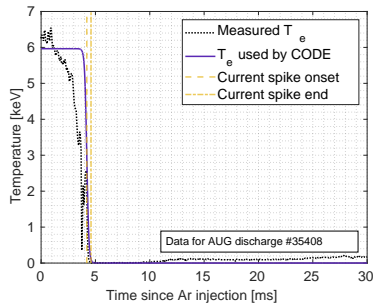


- Free electron temperature measured by ECE
- ECE blocked by high e^- densities on edge ≈ 1 ms after injection
- Assume exponential thermal quench



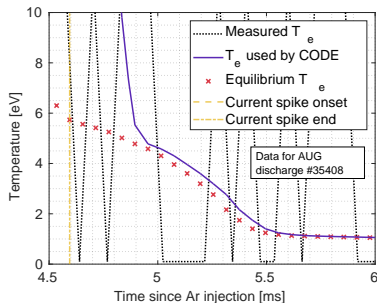
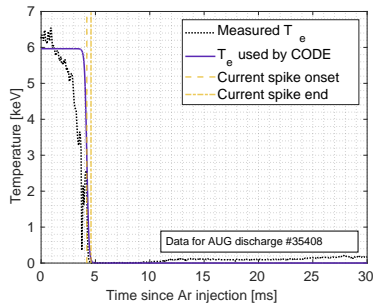


- Free electron temperature measured by ECE
- ECE blocked by high e^- densities on edge ≈ 1 ms after injection
- Assume exponential thermal quench
- But what is the final temperature?



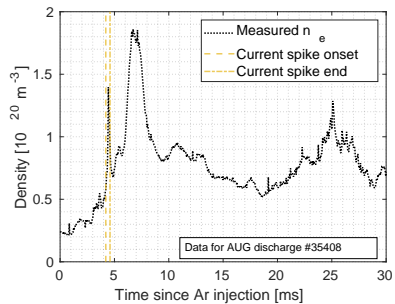
- Free electron temperature measured by ECE
- ECE blocked by high e^- densities on edge ≈ 1 ms after injection
- Assume exponential thermal quench
- But what is the final temperature?
- Best guess: Calculate assuming collisional-radiative equilibrium at prevailing D/Ar densities and current density

$$J^2 \sigma(T_e, Z_{\text{eff}}(T_e)) = \sum_i n_e(T_e) n_i L_i(T, n_e(T_e))$$

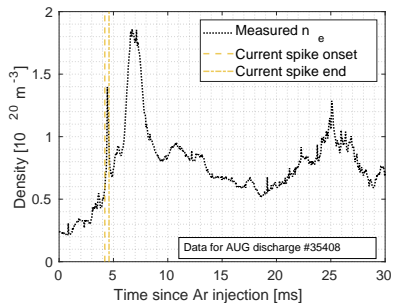


- Free electron temperature measured by ECE
- ECE blocked by high e^- densities on edge ≈ 1 ms after injection
- Assume exponential thermal quench
- But what is the final temperature?
- Best guess: Calculate assuming collisional-radiative equilibrium at prevailing D/Ar densities and current density

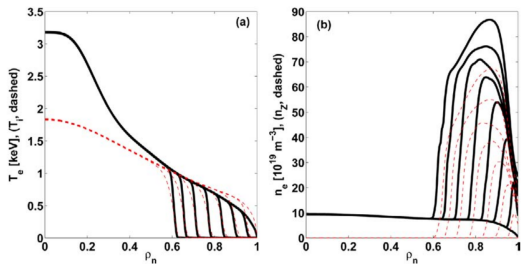
$$J^2 \sigma(T_e, Z_{\text{eff}}(T_e)) = \sum_i n_e(T_e) n_i L_i(T, n_e(T_e))$$



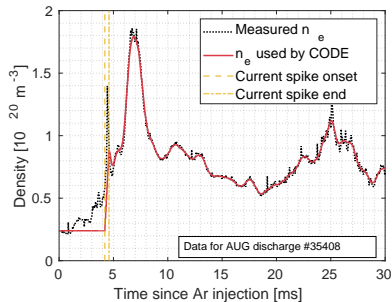
- Free electron density n_e measured by CO_2 interferometry



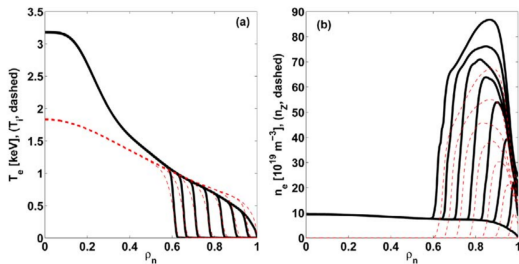
- Free electron density n_e measured by CO_2 interferometry
- Initial increase at the plasma edge

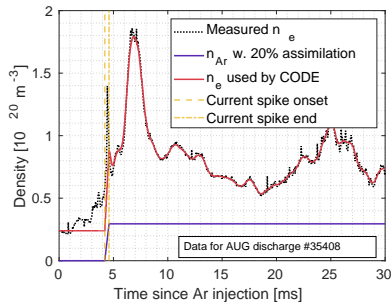


T_e and n_e profiles from [5] E. Fable, et al.: *Nuclear Fusion*, **56**, 2016

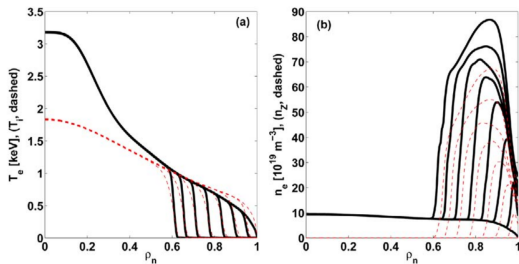


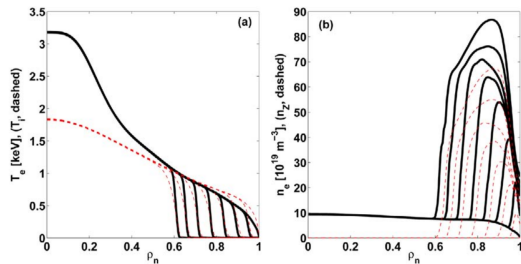
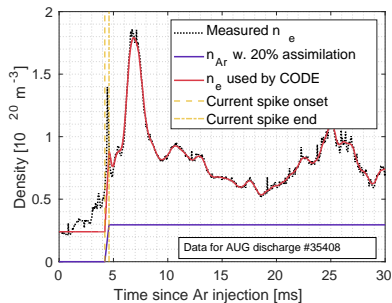
- Free electron density n_e measured by CO_2 interferometry
- Initial increase at the plasma edge
- Density on-axis constant until MHD mixing





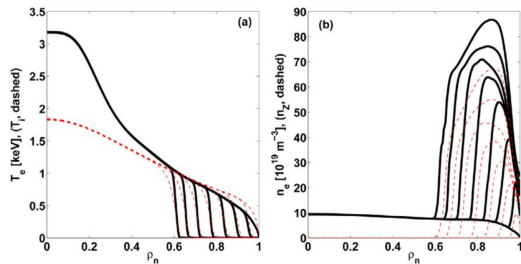
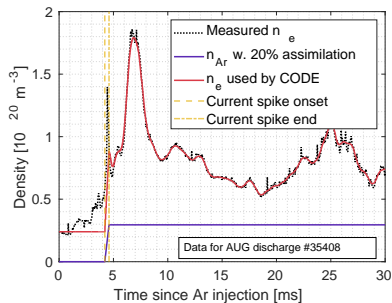
- Free electron density n_e measured by CO_2 interferometry
- Initial increase at the plasma edge
- Density on-axis constant until MHD mixing
- On-axis argon density assumed to evolve similarly





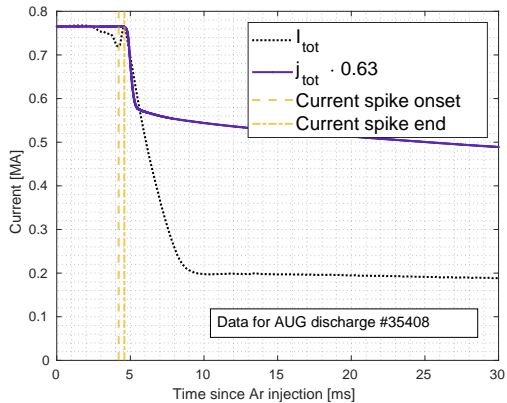
- Free electron density n_e measured by CO_2 interferometry
- Initial increase at the plasma edge
- Density on-axis constant until MHD mixing
- On-axis argon density assumed to evolve similarly
- Argon density after MHD mixing assumed constant, corresponding to homogenous distribution within pressure vessel (20% assimilation*)

T_e and n_e profiles from [5] E. Fable, et al.: *Nuclear Fusion*, **56**, 2016
 * 20% of total Ar in the plasma volume = 20% of vacuum vessel volume.

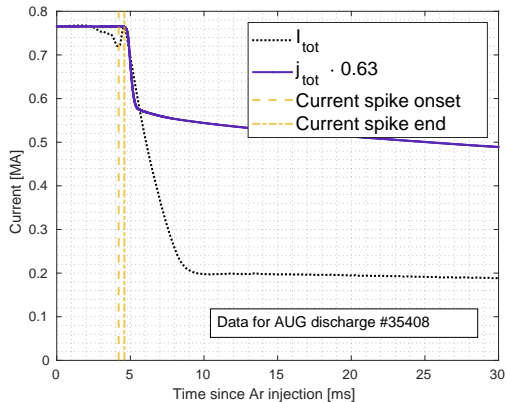


- Free electron density n_e measured by CO_2 interferometry
- Initial increase at the plasma edge
- Density on-axis constant until MHD mixing
- On-axis argon density assumed to evolve similarly
- Argon density after MHD mixing assumed constant, corresponding to homogenous distribution within pressure vessel (20% assimilation*)
- "Average" ionization state of argon chosen to give the measured n_e

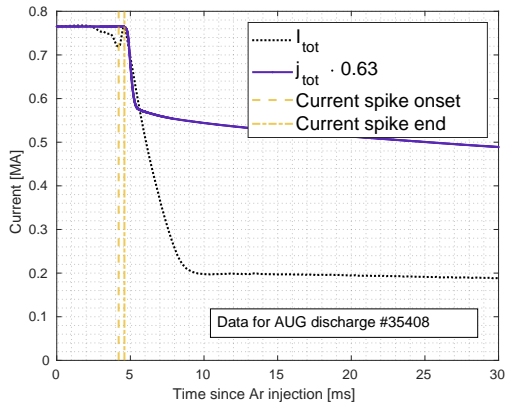
T_e and n_e profiles from [5] E. Fable, et al.: *Nuclear Fusion*, **56**, 2016
 * 20% of total Ar in the plasma volume = 20% of vacuum vessel volume.



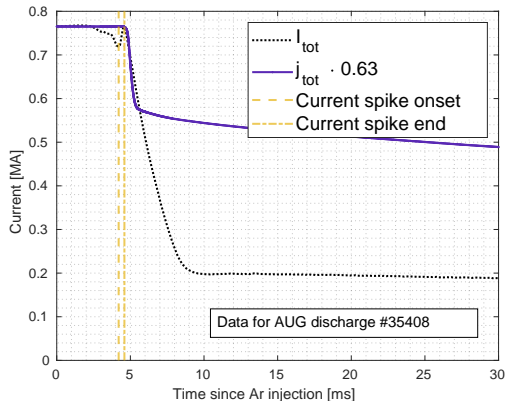
- CODE is 0D, so I and j are not directly comparable...



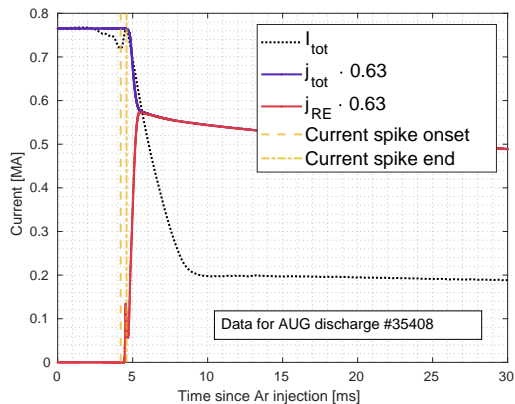
- CODE is 0D, so I and j are not directly comparable...
- Scaling factor 0.63 chosen to get $j_0 \approx 1.2 \text{ MA/m}^2$



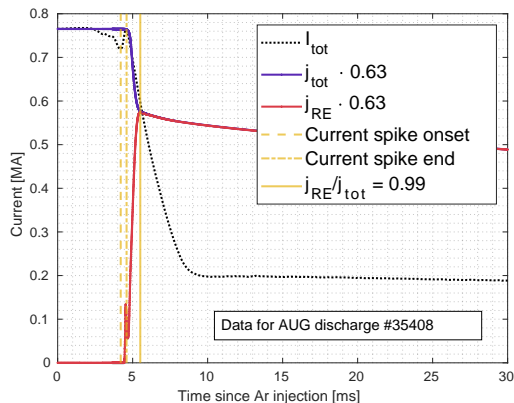
- CODE is 0D, so I and j are not directly comparable...
- Scaling factor 0.63 chosen to get $j_0 \approx 1.2 \text{ MA/m}^2$
- Current drops due to density/resistivity increase



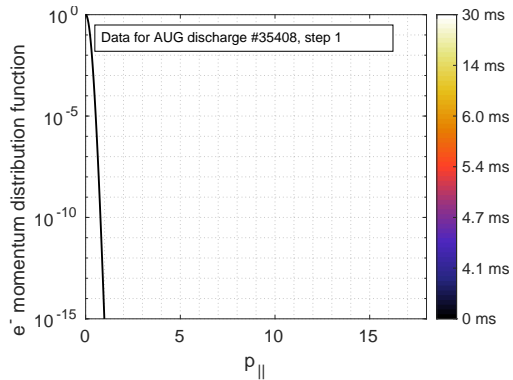
- CODE is 0D, so I and j are not directly comparable...
- Scaling factor 0.63 chosen to get $j_0 \approx 1.2 \text{ MA/m}^2$
- Current drops due to density/resistivity increase
- But why not to zero?



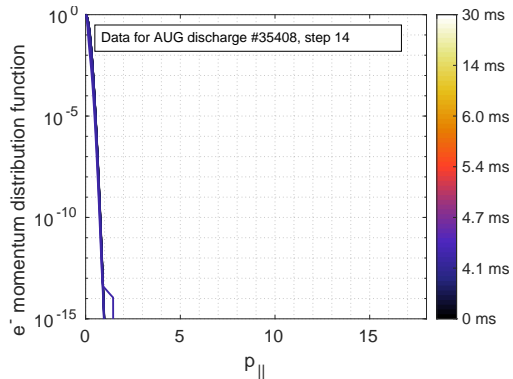
- CODE is 0D, so I and j are not directly comparable...
- Scaling factor 0.63 chosen to get $j_0 \approx 1.2 \text{ MA/m}^2$
- Current drops due to density/resistivity increase
- But why not to zero?
- The runaway current!



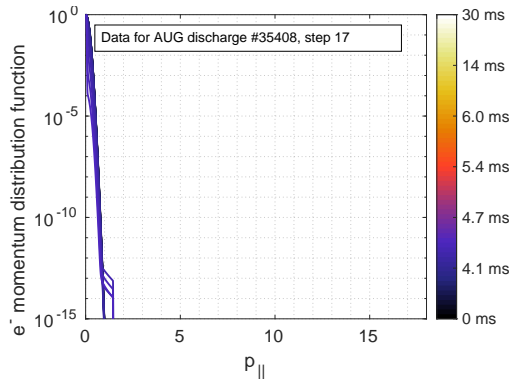
- CODE is 0D, so I and j are not directly comparable...
- Scaling factor 0.63 chosen to get $j_0 \approx 1.2 \text{ MA/m}^2$
- Current drops due to density/resistivity increase
- But why not to zero?
- The runaway current!
- Some ms after injection, 99% of the current is runaway!



■ Initially Maxwellian

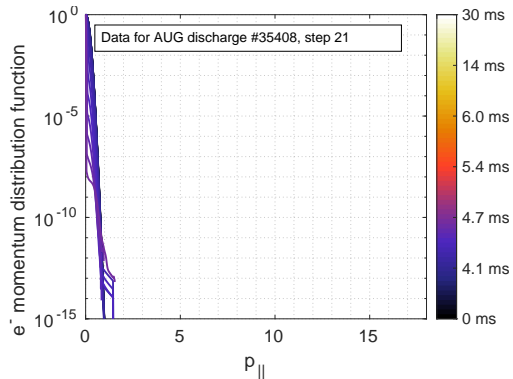


■ Initially Maxwellian

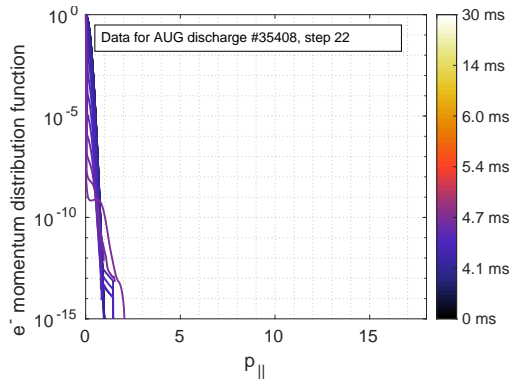


■ Initially Maxwellian

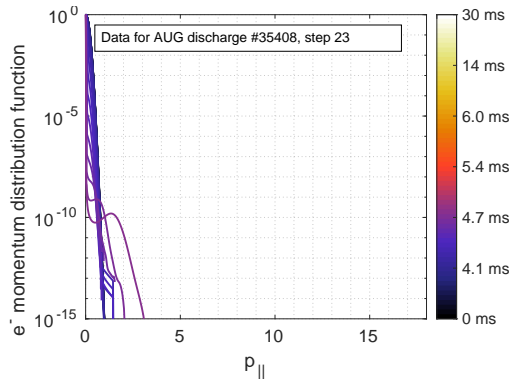
■ Maxwellian narrows as T decreases



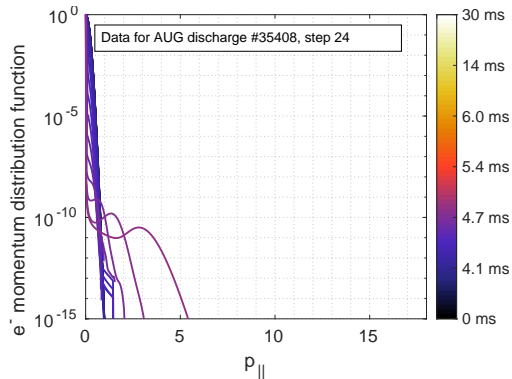
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed



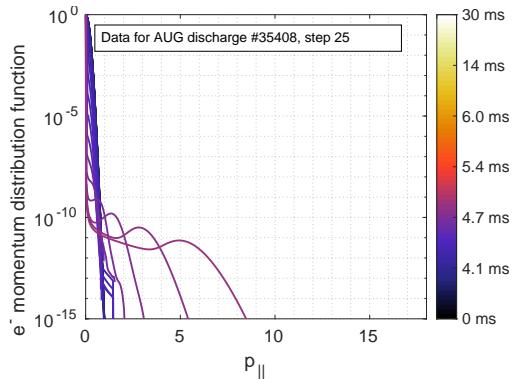
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed



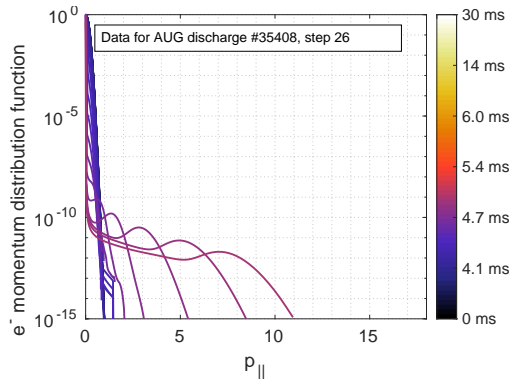
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



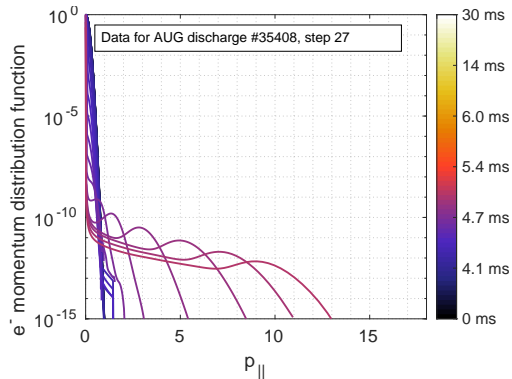
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



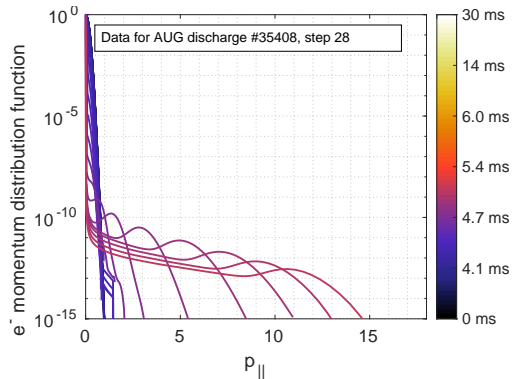
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



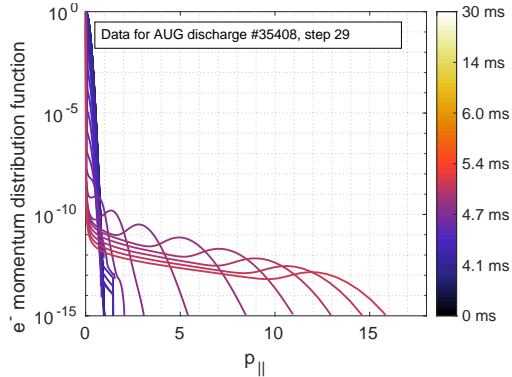
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



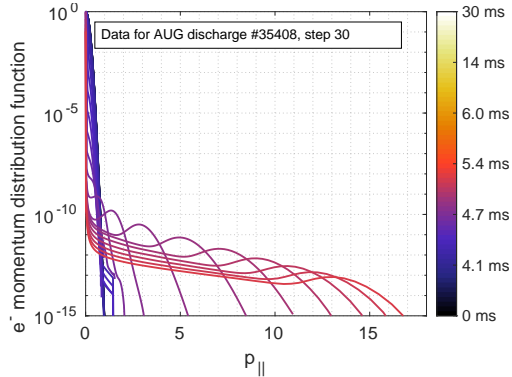
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



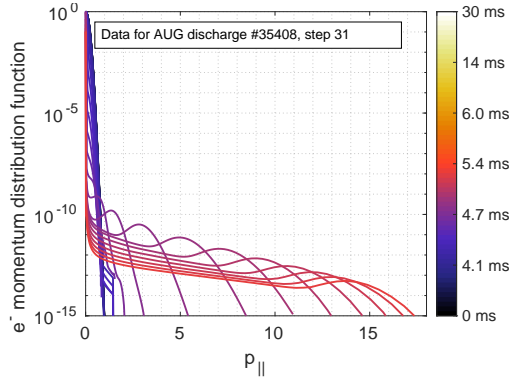
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



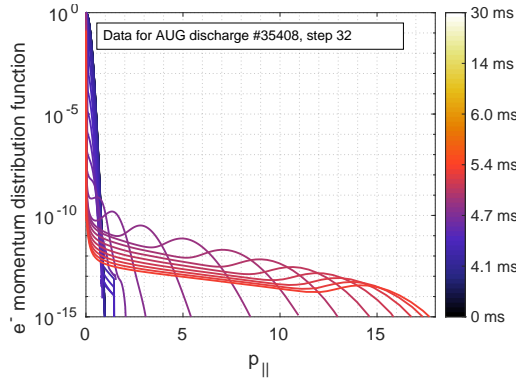
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



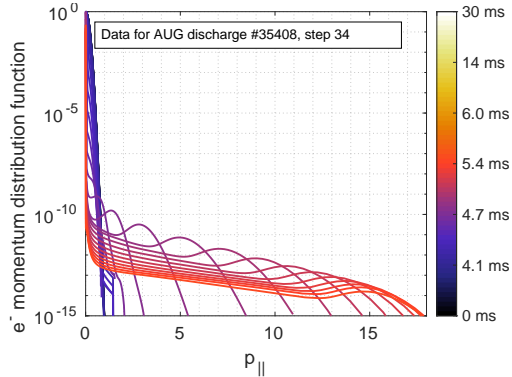
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



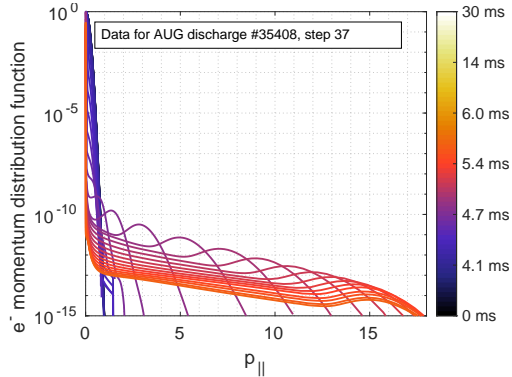
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



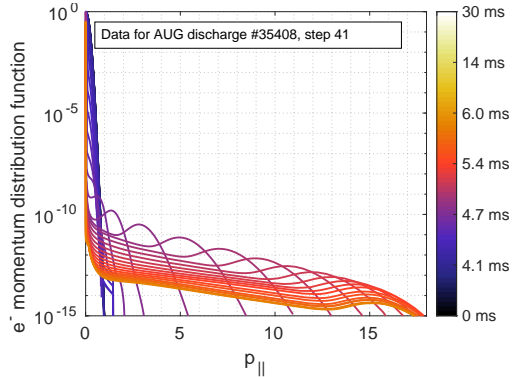
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



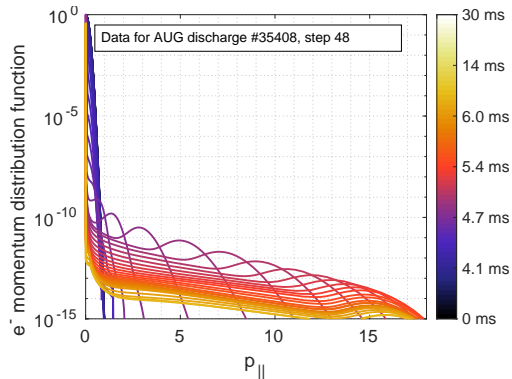
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



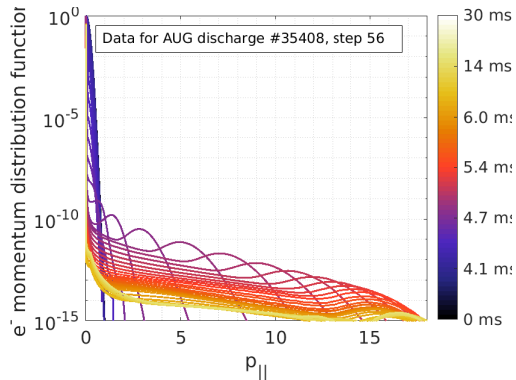
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



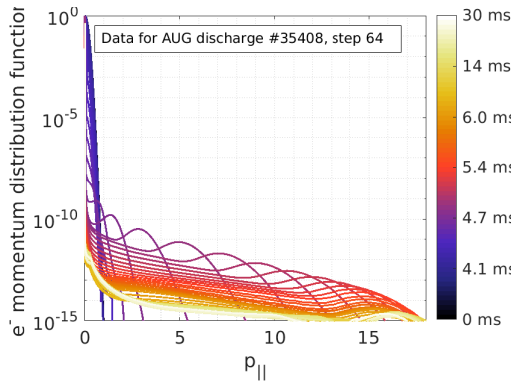
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



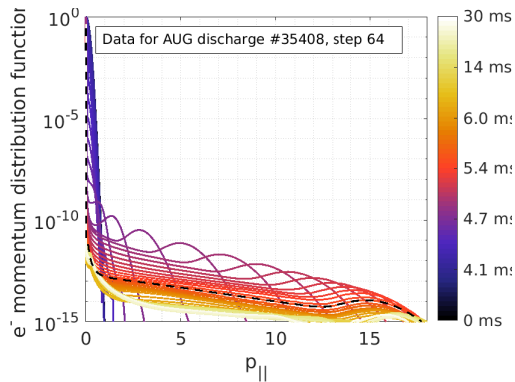
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



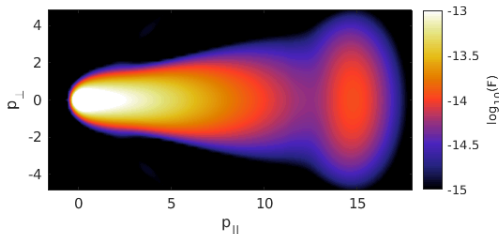
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated

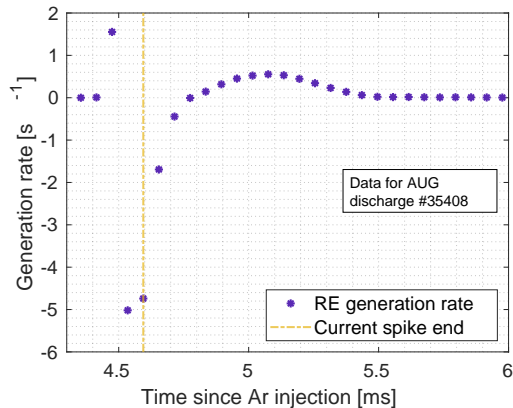


- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated



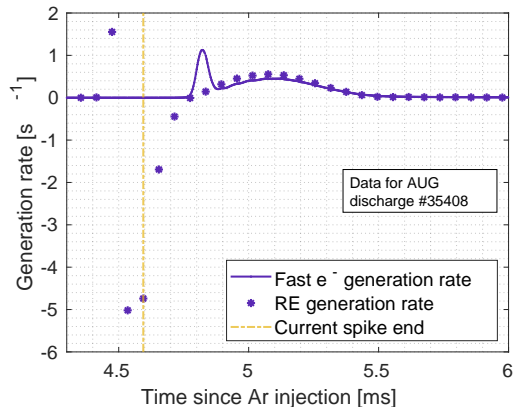
- Initially Maxwellian
- Maxwellian narrows as T decreases
- Hot-tail seed is formed
- Part of hot-tail seed accelerated
- 2D distribution at $J_{RE}/j_{tot} = 0.99$





- REs defined as e^- with $p > p_c$

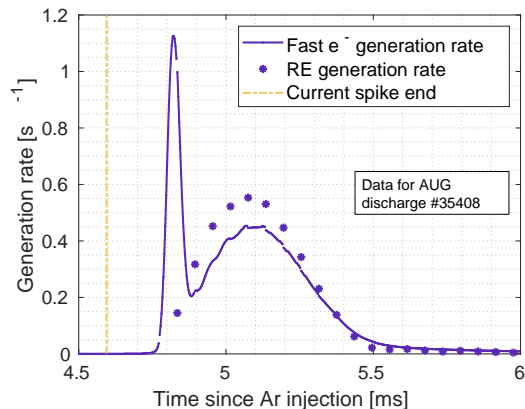
$$p_c = \frac{1}{\sqrt{E/E_c - 1}}, E_c = \frac{n_e e^3 \ln \Lambda}{4\pi \epsilon_0^2 m_e c^2}$$



- REs defined as e⁻ with $p > p_c$

$$p_c = \frac{1}{\sqrt{E/E_c - 1}}, E_c = \frac{n_e e^3 \ln \Lambda}{4\pi \epsilon_0^2 m_e c^2}$$

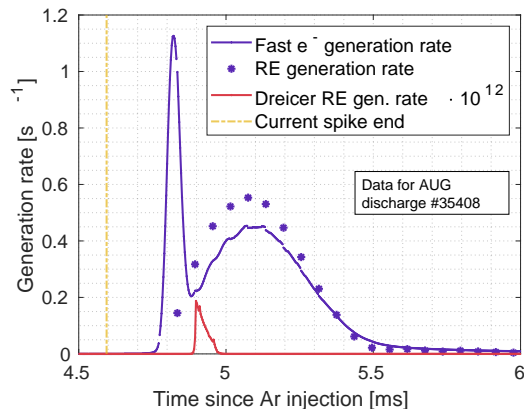
- Fast e⁻ defined as e⁻ with $p/mc > 0.75$



- REs defined as e⁻ with $p > p_c$

$$p_c = \frac{1}{\sqrt{E/E_c - 1}}, E_c = \frac{n_e e^3 \ln \Lambda}{4\pi \epsilon_0^2 m_e c^2}$$

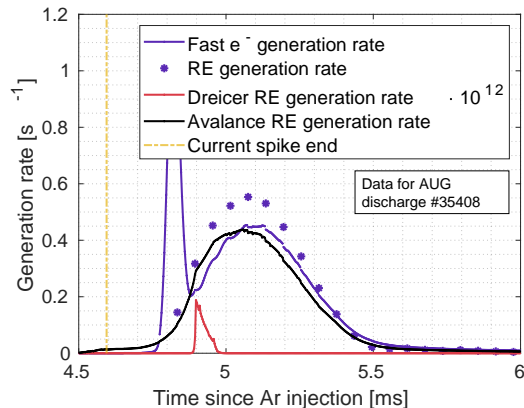
- Fast e⁻ defined as e⁻ with $p/mc > 0.75$



- REs defined as e^- with $p > p_c$

$$p_c = \frac{1}{\sqrt{E/E_c - 1}}, E_c = \frac{n_e e^3 \ln \Lambda}{4\pi \epsilon_0^2 m_e c^2}$$

- Fast e^- defined as e^- with $p/mc > 0.75$
- Dreicer generation rate calculated using neural network [2]



- REs defined as e⁻ with $p > p_c$

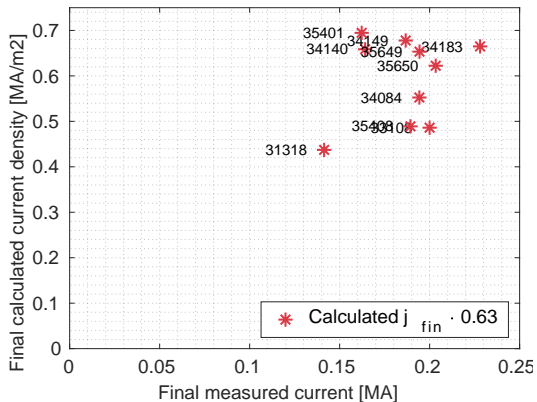
$$p_c = \frac{1}{\sqrt{E/E_c - 1}}, E_c = \frac{n_e e^3 \ln \Lambda}{4\pi \epsilon_0^2 m_e c^2}$$

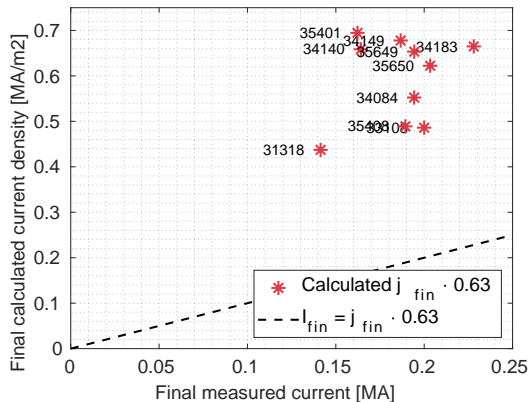
- Fast e⁻ defined as e⁻ with $p/mc > 0.75$
- Dreicer generation rate calculated using neural network [2]
- Avalanche growth rate calculated using semi-analytical formula [6]

[2] L. Hesslow, et al.: *Journal of Plasma Physics*, **85**(6), 2019

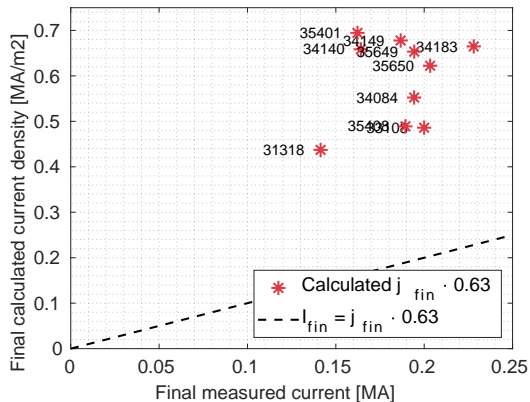
[6] L. Hesslow, et al.: *Nuclear Fusion*, **59**, 2019

- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$

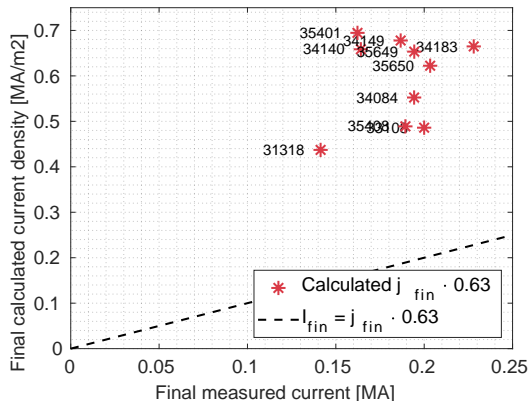




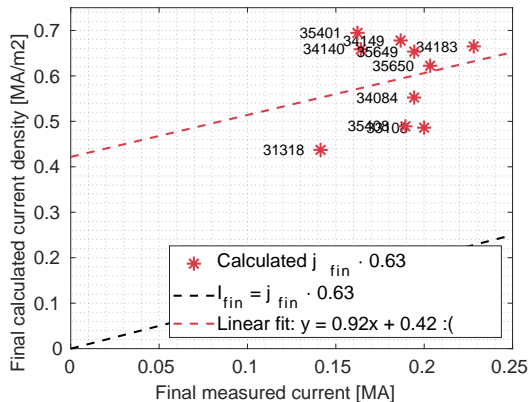
- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:



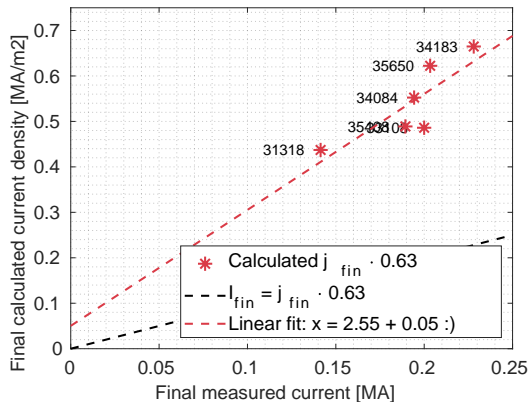
- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses



- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses
 - ▶ The current profile changes (0.63 scaling only valid pre-disruption)

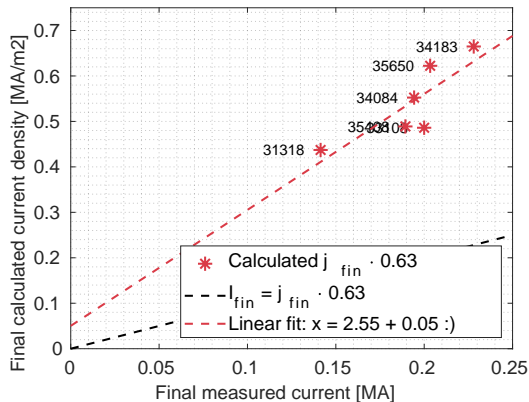


- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses
 - ▶ The current profile changes (0.63 scaling only valid pre-disruption)
- But we would have liked a line through origin and a better fit...



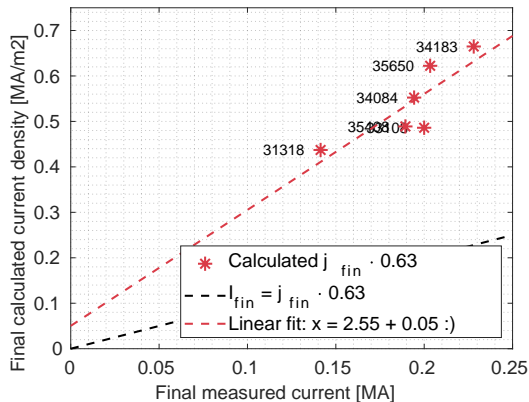
- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses
 - ▶ The current profile changes (0.63 scaling only valid pre-disruption)
- But we would have liked a line through origin and a better fit...
- ...which we get if we remove four "fishy" shots*!

* where the assumed exponential temperature decay falls to the final temperature without a plateau at a higher calculated equilibrium temperature.



- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses
 - ▶ The current profile changes (0.63 scaling only valid pre-disruption)
- But we would have liked a line through origin and a better fit...
- ...which we get if we remove four "fishy" shots*!
- Difference in slope expected (transport + profile change)

* where the assumed exponential temperature decay falls to the final temperature without a plateau at a higher calculated equilibrium temperature.



- Comparing the measured $I_{t=30ms}$ with calculated $j_{t=30ms} \cdot 0.63$
- $I_{t=30ms} \neq j_{t=30ms} \cdot 0.63$ since:
 - ▶ We don't model transport losses
 - ▶ The current profile changes (0.63 scaling only valid pre-disruption)
- But we would have liked a line through origin and a better fit...
- ...which we get if we remove four "fishy" shots*!
- Difference in slope expected (transport + profile change)
- To do: Improve temperature evolution estimate.

* where the assumed exponential temperature decay falls to the final temperature without a plateau at a higher calculated equilibrium temperature.

- Measured data from AUG shots was used (with some modification/interpretation):

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states

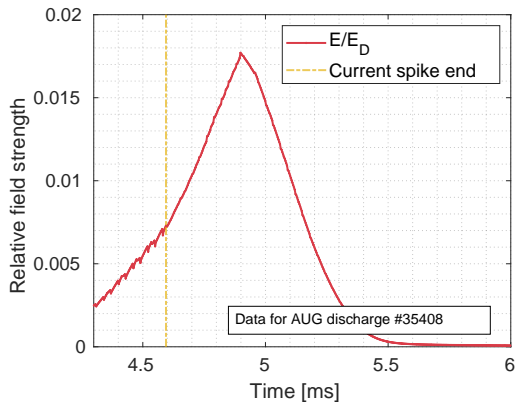
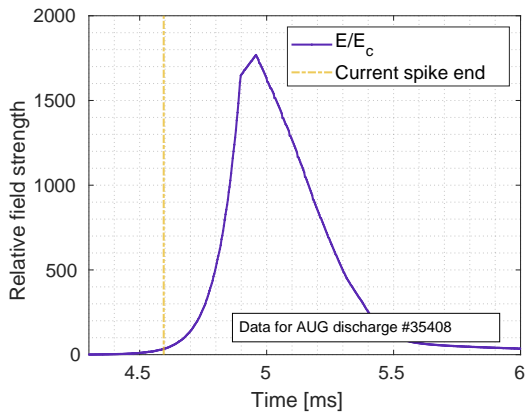
- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states
 - ▶ Initial on-axis current density

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states
 - ▶ Initial on-axis current density
- Momentum distribution and on-axis current density calculated with CODE

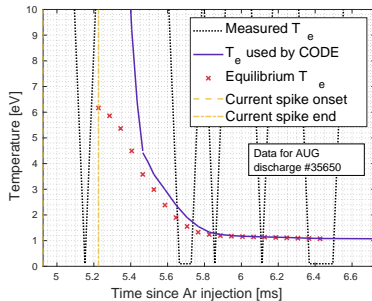
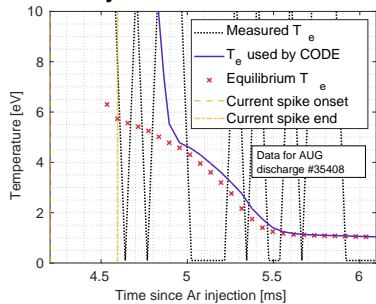
- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states
 - ▶ Initial on-axis current density
- Momentum distribution and on-axis current density calculated with CODE
- Hot-tail and avalanche most important RE generation mechanisms

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states
 - ▶ Initial on-axis current density
- Momentum distribution and on-axis current density calculated with CODE
- Hot-tail and avalanche most important RE generation mechanisms
- Calculated on-axis current density scales with measured current at $t = 30$ s

- Measured data from AUG shots was used (with some modification/interpretation):
 - ▶ On-axis free electron temperature
 - ▶ On-axis free electron density
- Some parameters were estimated/guessed:
 - ▶ On-axis argon density
 - ▶ On-axis argon ionization states
 - ▶ Initial on-axis current density
- Momentum distribution and on-axis current density calculated with CODE
- Hot-tail and avalanche most important RE generation mechanisms
- Calculated on-axis current density scales with measured current at $t = 30$ s
- Conclusion: CODE captures important features of RE generation



Non-fishy shots:



Fishy shots:

