





# Theory-supported analysis of runaway distributions using synchrotron radiation

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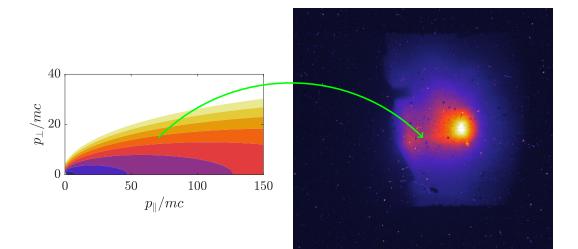
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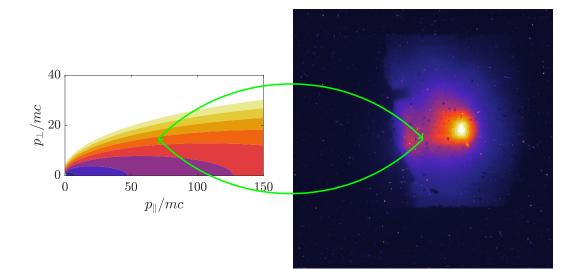
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- <sup>†</sup>See the author list of "B. Labit et al. 2019 Nucl. Fusion 59 086020"

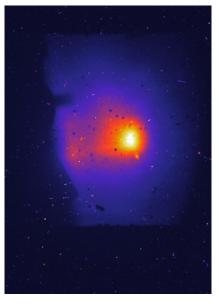


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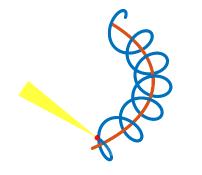




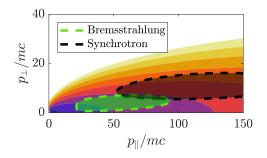


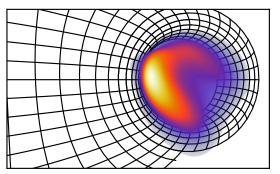


- 1. Synchrotron radiation
- 2. Soft
- 3. Theory-supported analysis in ASDEX-U
  - ► Fluid-kinetic simulations
  - Model to fit radial dynamics



- Emitted by **highly relativistic** particles...
  - ► (such as runaways)
  - ▶ ... in forward direction...
  - ▶ ... with continous spectrum...
- Typically observed at visible wavelengths in tokamaks
- Sensitive to details of  $f(r, p, \theta)$

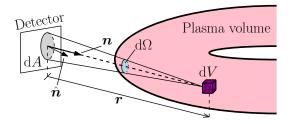




- Synthetic synchrotron (and bremsstrahlung) diagnostic
- Given  $f(r, p, \theta)$  and **magnetic field**, reproduces radiation pattern
- Applied to Alcator C-Mod, ASDEX Upgrade, DIII-D, JET and TCV
- Freely available on GitHub<sup>‡</sup>
- Python analysis framework under continuous development\*

<sup>†</sup> M. Hoppe et al, NF **58** *026032* (2018) <sup>‡</sup> https://github.com/hoppe93/S0FT2

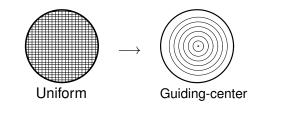
\* https://github.com/hoppe93/softplot



$$I = \iiint \Theta\left(\frac{\boldsymbol{r}}{r}\right) \frac{\boldsymbol{r} \cdot \hat{\boldsymbol{n}}}{r^3} \frac{\mathrm{d}I(\boldsymbol{x}, \boldsymbol{p}, \boldsymbol{r})}{\mathrm{d}\Omega} f(\boldsymbol{x}, \boldsymbol{p}) \,\mathrm{d}\boldsymbol{p} \mathrm{d}V \mathrm{d}A$$

- I = Radiation quantity
- f = Distribution function
- $rac{\mathrm{d}I}{\mathrm{d}\Omega} =$  Angular distribution of radiation

$$\Theta\left(rac{m{r}}{r}
ight) = egin{cases} 1, ext{if } m{r}/r ext{ is in the FOV} \ 0, ext{otherwise} \end{cases}$$



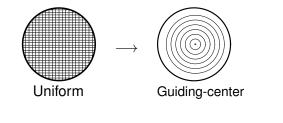
$$I = \int f(\rho, p_{\parallel}, p_{\perp}) \Theta\left(\frac{\boldsymbol{r}}{r}\right) \frac{\boldsymbol{r} \cdot \hat{\boldsymbol{n}}}{r^{3}} \frac{\mathrm{d}I}{\mathrm{d}\Omega} \underbrace{p_{\perp}J \,\mathrm{d}\rho \mathrm{d}\tau \mathrm{d}\phi}_{\mathrm{d}\boldsymbol{X}} \underbrace{\mathrm{d}p_{\parallel} \mathrm{d}p_{\perp} \mathrm{d}\zeta}_{\mathrm{d}\boldsymbol{p}} \,\mathrm{d}A.$$

#### where

- $\rho = {\rm Maximum}$  major radius along orbit
- $\tau = \mbox{Time}$  parameter along orbit

 $\phi = \text{Toroidal angle}$ 

 $\zeta =$  Gyro angle J = Trajectory Jacobian

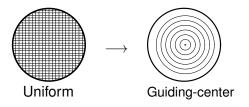


$$I = \int f(\rho, p_{\parallel}, p_{\perp}) \underbrace{\Theta\left(\frac{\boldsymbol{r}}{r}\right) \frac{\boldsymbol{r} \cdot \hat{\boldsymbol{n}}}{r^{3}} \frac{\mathrm{d}I}{\mathrm{d}\Omega} J \,\mathrm{d}\tau \mathrm{d}\phi \mathrm{d}\zeta \mathrm{d}A}_{\mathcal{G}(\rho, p_{\parallel}, p_{\perp})} p_{\perp} \mathrm{d}\rho \mathrm{d}p_{\parallel} \mathrm{d}p_{\perp}.$$

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$$I = \int f(\rho, p_{\parallel}, p_{\perp}) \mathcal{G}(\rho, p_{\parallel}, p_{\perp}) p_{\perp} d\rho dp_{\parallel} dp_{\perp}.$$

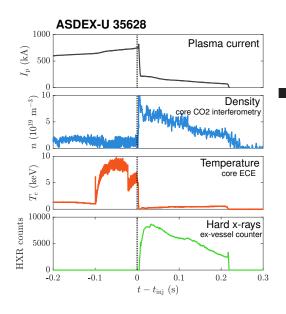
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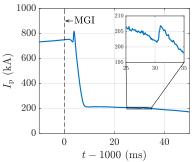
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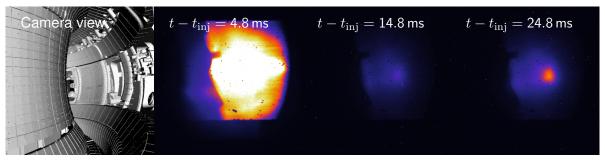
- $\zeta =$ Gyro angle
- J = Trajectory Jacobian

### **Experiment: ASDEX-U 35628**



- Massive Gas Injection (Ar),  $N_{\rm Ar} \approx 10^{21}$  particles
- $\blacktriangleright~$  Current:  $\sim 800\,\text{kA}$  to  $\sim 200\,\text{kA}$
- $\blacktriangleright~$  Temperature:  $\sim 5\,{\rm keV}$  to  $\sim 1\,{\rm eV}$
- ICRH applied
- Fast (1 kHz) visible-light camera
  - ► Equipped with 709 ± 9 nm filter (to remove line radiation)
  - Excellent video data!

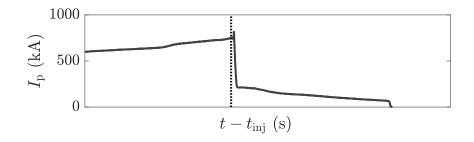


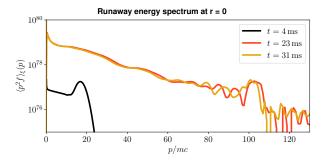


 $t - t_{
m inj} = 28.8\,
m ms$   $t - t_{
m inj} = 29.8\,
m ms$   $t - t_{
m inj} = 39.8\,
m ms$   $t - t_{
m inj} = 72.8\,
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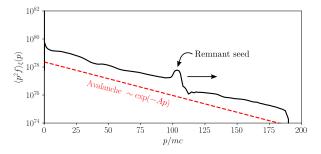
## **Theory-supported analysis**

- Fluid (GO) + Kinetic (CODE)
- Disruption ⇒ many experimental uncertainties
- Avalanche dynamics robust ⇒ try match final current (200 kA)
  - Skip TQ
  - Start with appropriately sized hot-tail seed
  - Evolve through CQ and plateau

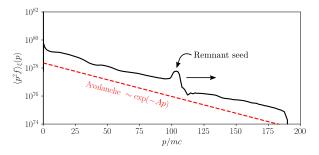




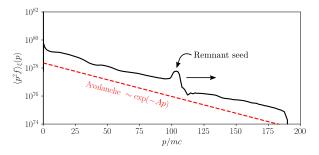
Seed is accelerated ( $\tau_{\rm acc} \lesssim 5 \, {\rm ms}$ )



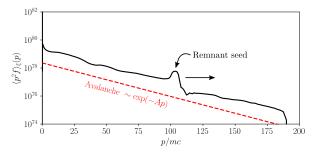
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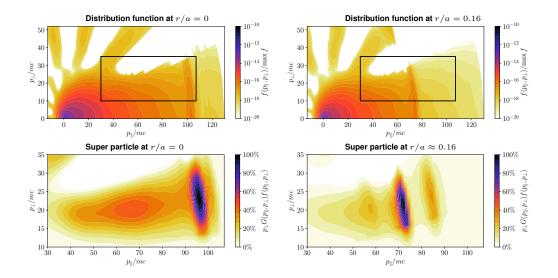
- Seed is accelerated ( $\tau_{\rm acc} \lesssim 5 \, {\rm ms}$ )
- Gradual formation of avalanche distribution

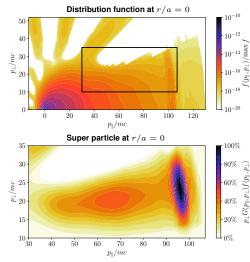


- Seed is accelerated ( $\tau_{\rm acc} \lesssim 5 \, {\rm ms}$ )
- Gradual formation of avalanche distribution
- $\blacksquare \quad \text{Spot shape change after} \sim 27 \ \text{ms}$

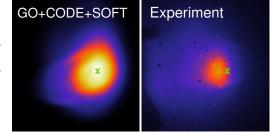


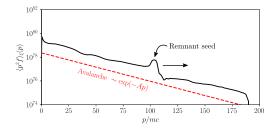
- Seed is accelerated ( $\tau_{\rm acc} \lesssim 5 \, {\rm ms}$ )
- Gradual formation of avalanche distribution
- Spot shape change after ~ 27 ms
- Slow avalanching ( $\tau_{ava} \gg 27 \text{ ms}$ )





- Synchrotron dominated by seed
- Seed energy varies with radius
- Spot shape type matches but too large





Recall: Only dominant emitters matter

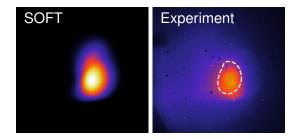
 $\implies$  in a seed dominated scenario, GO+CODE simulations suggest

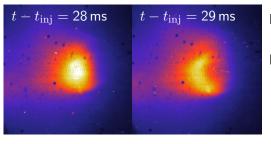
$$f(r, p, \xi) = f_r(r)\delta(p - p^*)\exp(C\xi)$$

- Synchrotron simulations insensitive to exact  $p^*$
- Pitch distribution is (approximately) steady state

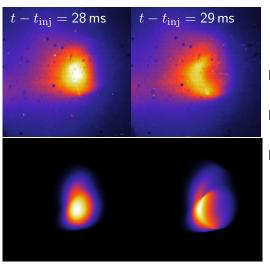
$$f(r, p, \xi) = f_r(r)\delta(p - p^*)\exp(C\xi)$$

Good fit: 
$$\begin{cases} f_r(r) & \propto 1 - r/(17 \text{ cm}), \\ p^* &= 42mc, \\ C &= 100, \end{cases}$$

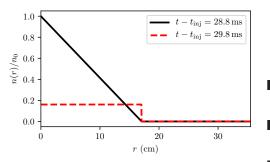


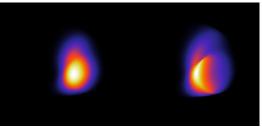


- Synchrotron spot transition, ellipse  $\rightarrow$  crescent
- Sub-ms event: assume momentum distribution changes negligibly



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- Consistent with runaway profile flattening (reconnection)





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- Sub-ms event: assume momentum distribution changes negligibly
- ⇒ Consistent with runaway profile flattening (reconnection)

- GO+CODE+SOFT simulations of ASDEX-U disruption suggest we observe the **runaway seed** on visible camera
- Motivates simple model for synchrotron fitting purposes
- Synchrotron spot shape change in ASDEX-U #35628 consistent with flattened runaway density profile